Orbital Angular Momentum

Now we have that $\hat{r}=\sqrt{\sum_{j} \hat{x}_{j}{ }^{2}}$. We also have that $\left[\hat{L}_{i}, \sum_{j} \hat{x}_{j}^{2}\right]=0$. One can show from this that $\left[\hat{L}_{i}, \hat{r}\right]=0$.

$$
\left[L_{i}, f(\hat{r})\right]=\left[L_{i}, r\right] f^{\prime}(\hat{r})=0
$$

for amp function $f$ of $r$.

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\end{equation*}
$$

## Orbital Angular Momentum

So, for any spherically symmetric potential $V(r)$, we have that

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\left[\hat{L}_{i}, \hat{H}\right]=\left[\hat{L}_{i},-\frac{\hbar^{2}}{2 M} \nabla^{2}+V(r)\right]
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In other words, $\hat{H}, \hat{L}_{i}$ and $\hat{L}^{2}$ all commute with one another.
This is an important and powerful result.

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In other words, $\hat{H}, \hat{L}_{i}$ and $\hat{L}^{2}$ all commute with one another.
This is an important and powerful result. Given any 3D quantum system, we can find a basis of simultaneous eigenfunctions of $\hat{H}, \hat{L}^{2}$ and $\hat{L}_{3}$.

## Orbital Angular Momentum

We can translate the definitions of $\hat{L}_{i}$ to spherical polars. We have

$$
x_{1}=r \sin \theta \cos \phi, \quad x_{2}=r \sin \theta \sin \phi, \quad x_{3}=r \cos \theta
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\frac{\partial}{\partial \theta}=\sum_{i} \frac{\partial x_{i}}{\partial \theta} \frac{\partial}{\partial x_{i}}
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Thus

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\begin{aligned}
\frac{\partial}{\partial \theta} & =\sum_{i} \frac{\partial x_{i}}{\partial \theta} \frac{\partial}{\partial x_{i}} \\
& =r \cos \theta \cos \phi \frac{\partial}{\partial x_{1}}+r \cos \theta \sin \phi \frac{\partial}{\partial x_{2}}-r \sin \theta \frac{\partial}{\partial x_{3}}(7.52)
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\frac{\partial}{\partial \phi} & =\sum_{i} \frac{\partial x_{i}}{\partial \phi} \frac{\partial}{\partial x_{i}} \\
& =-r \sin \theta \sin \phi \frac{\partial}{\partial x_{1}}+r \sin \theta \cos \phi \frac{\partial}{\partial x_{2}} \tag{7.53}
\end{align*}
$$

## Orbital Angular Momentum

We thus obtain

$$
i \hbar\left(\cos \phi \cot \theta \frac{\partial}{\partial \phi}+\sin \phi \frac{\partial}{\partial \theta}\right)=-i \hbar\left(x_{2} \frac{\partial}{\partial x_{3}}-x_{3} \frac{\partial}{\partial x_{2}}\right)
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\end{align*}
$$

## Orbital Angular Momentum

We can also obtain

$$
\hat{L}^{2}=\sum_{i} \hat{L}_{i}^{2}=\left(\hat{L}_{1}+i \hat{L}_{2}\right)\left(\hat{L}_{1}-i \hat{L}_{2}\right)+i\left[\hat{L}_{1}, \hat{L}_{2}\right]+\hat{L}_{3}^{2}
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& =-\hbar^{2}\left(\cot \theta e^{i \phi} \frac{\partial}{\partial \phi}-i e^{i \phi} \frac{\partial}{\partial \theta}\right)\left(\cot \theta e^{-i \phi} \frac{\partial}{\partial \phi}+i e^{-i \phi} \frac{\partial}{\partial \theta}\right)
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& \quad+i \hbar^{2} \frac{\partial}{\partial \phi}-\hbar^{2} \frac{\partial^{2}}{\partial \phi^{2}} \\
& =-\hbar^{2}\left(\frac{\partial^{2}}{\partial \theta^{2}}+\cot \theta \frac{\partial}{\partial \theta}+\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \phi^{2}}\right)
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= & -\hbar^{2}\left(\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial}{\partial \theta}\right)+\frac{1}{\sin ^{2} \theta} \frac{\partial^{2}}{\partial \phi^{2}}\right) . \tag{7.57}
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Recall that $\left[\hat{L}^{2}, \hat{L}_{3}\right]=0$.

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We can thus seek simultaneous eigenfunctions of the form $Y(\theta) \exp (i m \phi)$, since $\hat{L}_{3} \exp (i m \phi)=\hbar m \exp (i m \phi)$. As $\phi$ is defined modulo $2 \pi$, we need $e^{i m(\phi+2 \pi)}=e^{i m \phi}$, so $e^{i 2 m \pi}=1$ and $m$ is an integer.

## Orbital Angular Momentum

This leaves us with an eigenvalue equation for $\hat{L}^{2}$ :

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-\hbar^{2}\left(\frac{1}{\sin \theta} \frac{d}{d \theta}\left(\sin \theta \frac{d}{d \theta}\right)-\frac{m^{2}}{\sin ^{2} \theta}\right) Y(\theta)=\lambda Y(\theta) \tag{7.59}
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From a physics perspective, the key fact about this equation is that we can show it has non-singular solutions if and only if $\lambda=\hbar^{2} I(I+1)$ for some integers $I \geq 0$

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From a physics perspective, the key fact about this equation is that we can show it has non-singular solutions if and only if $\lambda=\hbar^{2} I(I+1)$ for some integers $I \geq 0$ and for $m$ in the range $-I \leq m \leq I$.


## Orbital Angular Momentum

The solutions are called the associated Legendre functions $P_{l, m}(\theta)$.

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$$
\begin{equation*}
-\hbar^{2} \frac{d}{d w}\left(\left(1-w^{2}\right) \frac{d Y}{d w}\right)-\left(\lambda-\frac{m^{2}}{1-w^{2}}\right) Y=0 \tag{7.60}
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$$

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For $m=0$ and $\lambda=\hbar^{2} l(I+1)$ this is Legendre's differential equation for functions of degree $l$, which has solution $P_{l}(w)$.

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For $m=0$ and $\lambda=\hbar^{2} l(I+1)$ this is Legendre's differential equation for functions of degree $I$, which has solution $P_{l}(w)$. For general $m$ it's an associated Legendre differential equation.

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For $m=0$ and $\lambda=\hbar^{2} l(I+1)$ this is Legendre's differential equation for functions of degree $l$, which has solution $P_{l}(w)$. For general $m$ it's an associated Legendre differential equation.
The associated Legendre functions can be obtained from the Legendre polynomials $P_{I}$ by

$$
\begin{equation*}
P_{l, m}(\theta)=(\sin \theta)^{|m|} \frac{d^{|m|}}{d(\cos \theta)^{|m|}} P_{l}(\cos \theta) \tag{7.61}
\end{equation*}
$$

(up to normalisation. Note that the solutions for -m are proportional to those for m , for given I.)

## Orbital Angular Momentum

We thus have the overall solution given by the spherical harmonic with total angular momentum quantum number I and $\hat{L}_{3}$ quantum number $m$ :

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Y_{l, m}(\theta, \phi)=P_{l, m}(\theta) \exp (i m \phi),
$$

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$$
Y_{l, m}(\theta, \phi)=P_{l, m}(\theta) \exp (i m \phi)
$$

an eigenfunction of $\hat{L}^{2}$ and $\hat{L}_{3}$ with eigenvalues $\hbar^{2} I(I+1)$ and $\hbar m$ respectively.

## Orbital Angular Momentum

We thus have the overall solution given by the spherical harmonic with total angular momentum quantum number I and $\hat{L}_{3}$ quantum number $m$ :

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Y_{l, m}(\theta, \phi)=P_{l, m}(\theta) \exp (i m \phi)
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an eigenfunction of $\hat{L}^{2}$ and $\hat{L}_{3}$ with eigenvalues $\hbar^{2} I(I+1)$ and $\hbar m$ respectively.

For plots of some spherical harmonics see e.g. mathworld.wolfram.com/SphericalHarmonic.html.

## Solving the 3D Schrödinger equation for a spherically symmetric potential

The time-independent SE is

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\begin{equation*}
-\frac{\hbar^{2}}{2 M} \nabla^{2} \psi+V(r) \psi=E \psi \tag{7.62}
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Recall that in spherical polar coordinates

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\nabla^{2}=\frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} \frac{\partial}{\partial r}\right)+\frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial}{\partial \theta}\right)+\frac{1}{r^{2} \sin ^{2} \theta} \frac{\partial^{2}}{\partial \phi^{2}}
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So, we have a standard 1D radial Schrödinger equation for $\psi(r)$, with the modified potential $V(r)+\frac{\hbar^{2} /(l+1)}{2 M r^{2}}$.

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This makes sense physically, since a state $\psi$ with $\langle\underline{L}\rangle_{\psi} \neq \underline{0}$ by definition has a nonzero vector associated with it, which breaks spherical symmetry.


## Degeneracies

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## Solving the 3D Schrödinger equation for a spherically symmetric potential



## Theorem

The ground state (i.e. lowest energy bound state) solution of the 3D Schrödinger equation for a spherically symmetric potential must have $I=m=0$ and is thus spherically symmetric.

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I.e., the eigenstates $\psi_{i}$ are all spherically symmetric solutions. We can thus write $\psi(r)=\sum_{i} c_{i} \psi_{i}(r)$ for some constants $c_{i}$ such that $\sum_{i}\left|c_{i}\right|^{2}=1$.

Solving the 3D Schrödinger equation for a spherically symmetric potential

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E=\int_{r=0}^{\infty} \psi^{*}(r)\left(-\frac{\hbar^{2}}{2 M}\left(\frac{d^{2}}{d r^{2}}+\frac{2}{r} \frac{d}{d r}\right) \psi(r)\right.
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& \left.+\int_{r=0}^{\infty} \psi^{*}(r)\left(\frac{\hbar^{2}}{2 M r^{2}} I(I+1)\right) \psi(r)\right) \\
& >\int_{r=0}^{\infty} \psi^{*}(r)\left(-\frac{\hbar^{2}}{2 M}\left(\frac{d^{2}}{d r^{2}}+\frac{2}{r} \frac{d}{d r}+V(r)\right)\right) \psi(r)  \tag{7.71}\\
& =\left[\left|c_{j}\right|^{2} E_{j} \left\lvert\,=\int_{r=0}^{\infty} \sum_{i} c_{i} \psi_{i}^{*}(r)\left(-\frac{\hbar^{2}}{2 M}\left(\frac{d^{2}}{d r^{2}}+\frac{2}{r} \frac{d}{d r}+V(r)\right)\right) \sum_{j} c_{j} \psi_{j}(r)\right.\right. \text {. }
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Now this last term is

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We now have

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E=-\frac{M e^{4}}{32 \pi^{2} \epsilon_{0}^{2} \hbar^{2}} \frac{1}{N^{2}} .
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In fact, the true degeneracy of the Nth energy level of the hydrogen atom in a full non-relativistic quantum mechanical treatment is $2 N^{2}$ : the extra factor of 2 arises from an intrinsically quantum mechanical degree of freedom, the electron spin, which has no direct classical analogue.

## Towards the periodic table

We could try to generalize this discussion to atoms other than hydrogen. These have a nucleus with charge $+Z e$, orbited by $Z$ independent electrons, where the atomic number $Z$ is an integer greater than one.

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If we ignore this temporarily, we can obtain solutions of the form

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E=\sum_{i=1}^{Z} E_{i} \tag{7.79}
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So the lowest overall energy state is given by filling up the energy levels in order of increasing energy, starting with the lowest.

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"good" bomb will explode if photon hits trigger
"due" Lamb will not explode t the photon just reflect.

Mar problem. Yo r have a stuck of good t dud bombs, and $3^{d}$ lat how which is wish. want to identity soul good bombs.

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